



Remote Sensing Of Titan

- Lifting The Veil On The Origins Of Life?

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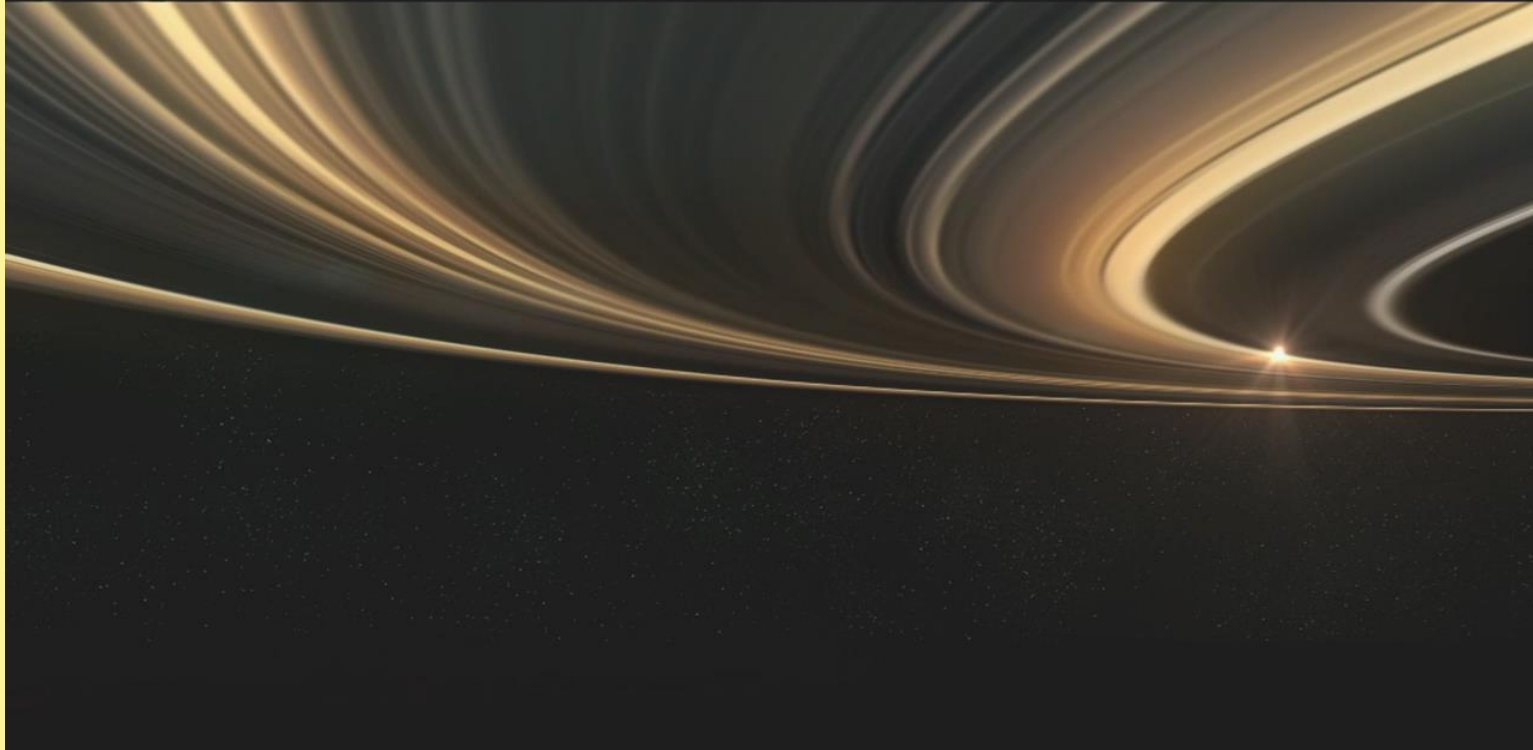
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Titan Was Designated A PRIME OBJECTIVE For Voyager Mission



BBC



Carl Sagan
(voiceover);
Astronomer,
Planetary
Scientist,
Astrophysicist
and Cosmologist
John Casani;
Voyager Project
Manager
Jim Bell;
Author of
"Interstellar Age"

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- When Voyager reached Titan in 1980 its cameras were defeated by 'smog'...

... BUT, Titan's Atmosphere Was Impenetrable To Voyager 1's Cameras...



BBC



Jim Bell;
Author of
"Interstellar Age"
Larry
Soderblom;
Voyager Imaging
Science Team
(Astrogeologist)

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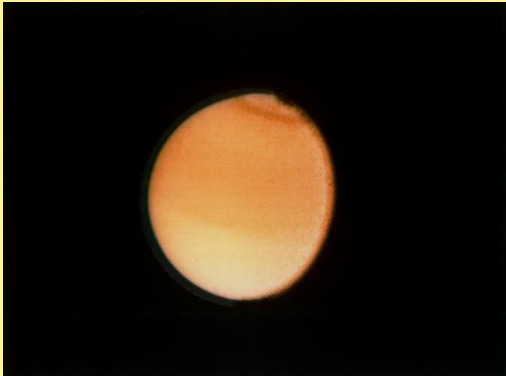
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- *Eventually, NASA/ESA/ISA plot collaborative return to Saturnian System*
 - *Cassini-Huygens orbiter and probe combination blasted off in 1997*
→ Aiming to penetrate 'the veil' of Titan's smog of 'THOLIN' organics...

Cassini-Huygens Mission Aims To Draw Back The Veil on Titan

- At Voyagers' flyby, Titan's haze had preserved that moon's mystery

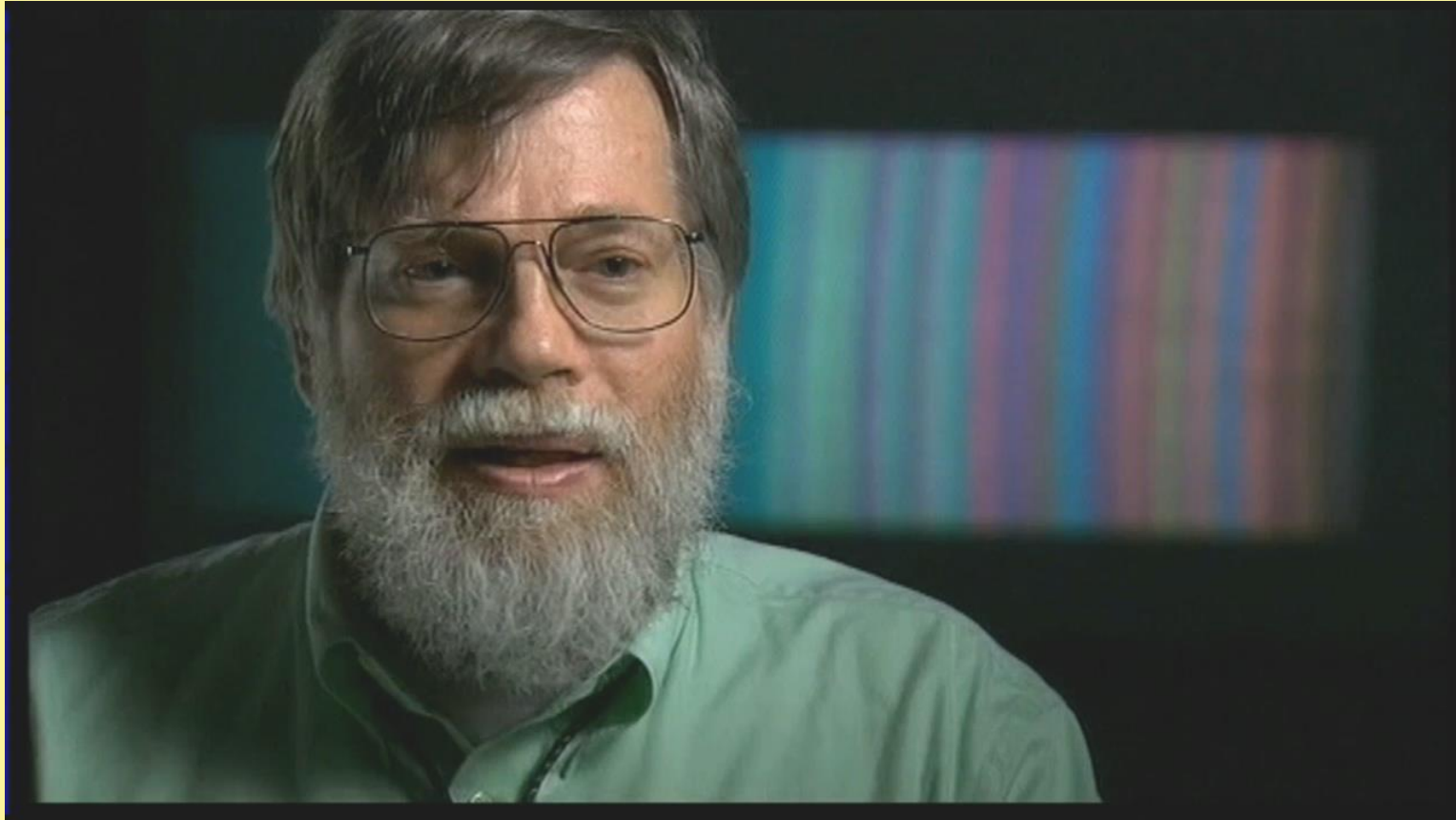


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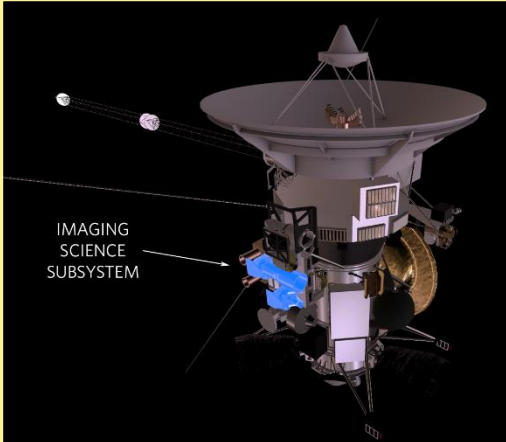
After The Voyager Imaging Disappointment, Cassini-Huygen's Instrument Package Was Carefully Considered For The Study of Titan...



Larry Esposito;
Principal
Investigator,
Cassini
Ultraviolet
Imaging
Spectrometer
(UVIS) Jack
Fortune;
Narrator

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*ISS Has Twin 1kx1k CCD Cameras Sensitive From Near-UV To Near-IR,
Including 938 nm, Which Confers Some Smog Penetration Ability...*

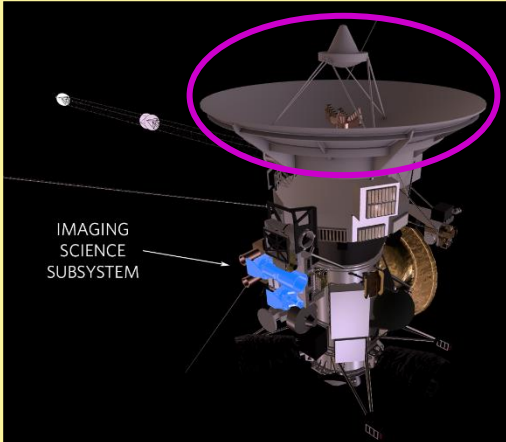


Jack Fortune;
Narrator
Caroline
Porco;
Team Leader,
Cassini Imaging
Science
Subsystem (ISS)

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The Active-Sensing, Synthetic Aperture Radar (Ku-Band; $\lambda = 2.18$ cm), Could Penetrate Thick Atmospheres With Ease...

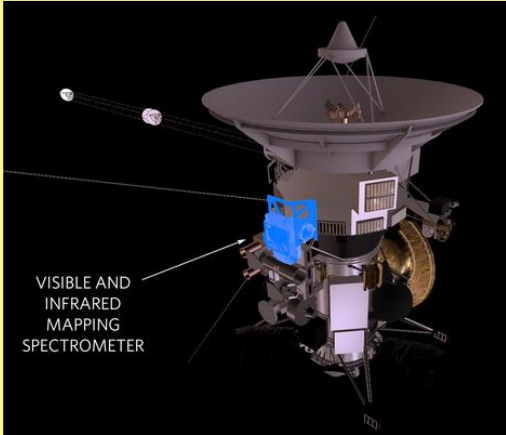


Jack Fortune;
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VIMS Scans From 0.35 μm (Near-UV) To 5.1 μm (Mid-IR) In 352 Bands & Delivers A 2-D Map Of Spectral Responses...

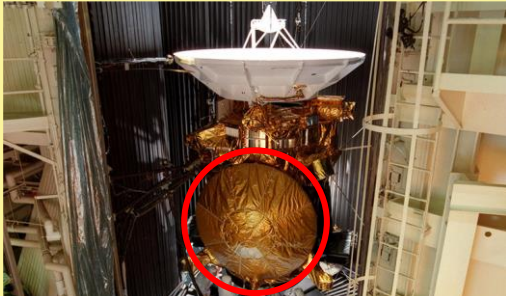


Jack Fortune;
Narrator
Bob Brown;
Team Leader,
Cassini Visual
and Infrared
Mapping
Spectrometer
(VIMS)

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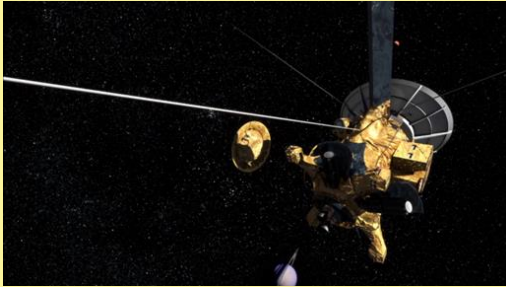
ESA's Huygens Probe Was Designed To Enter Titan's Atmosphere And Make Measurements All The Way Down To The Surface...



Jack Fortune;
Narrator
John Zarnecki;
Principal
Investigator,
Huygens Surface
Science Package
(SSP)

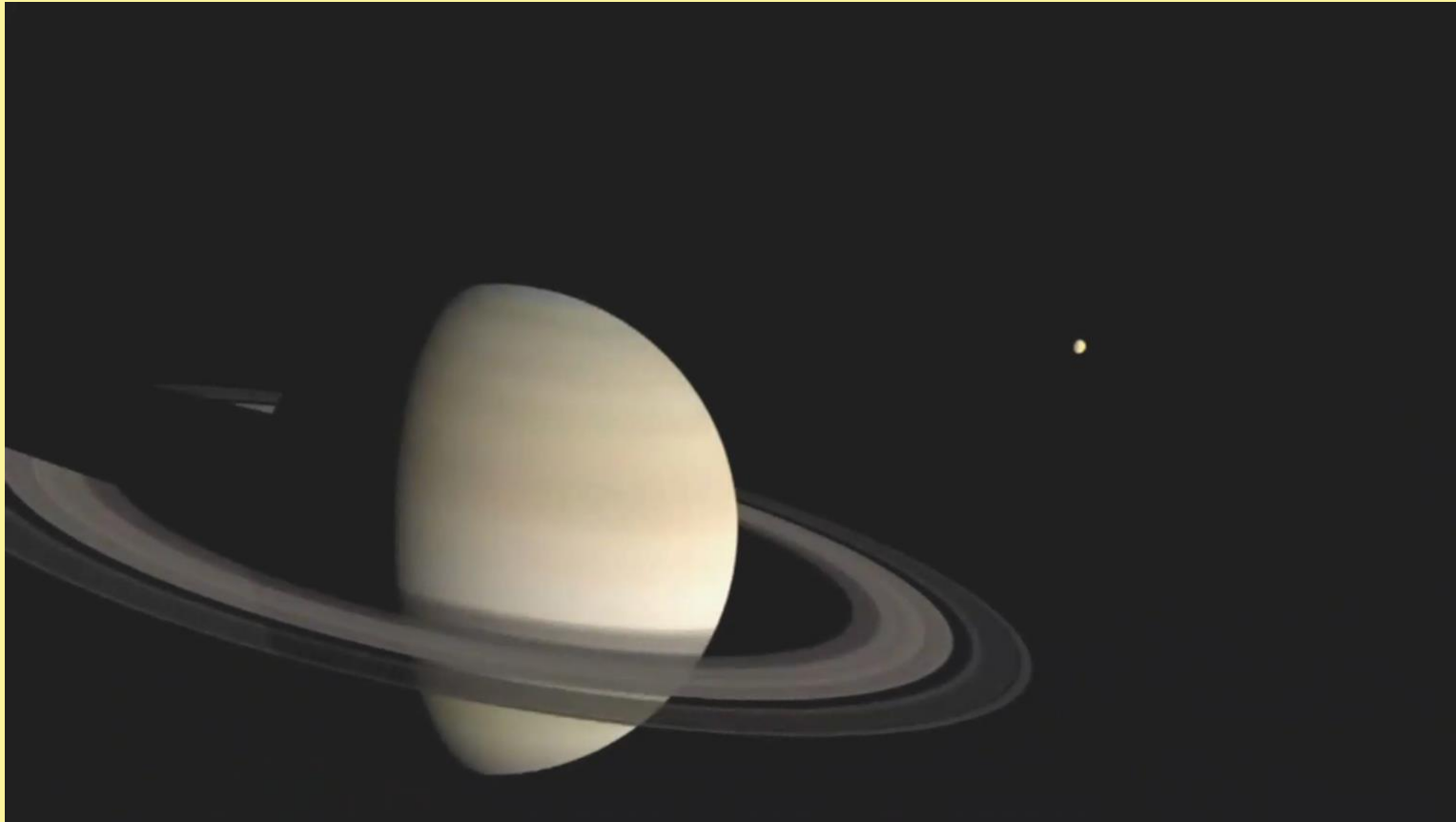
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On Christmas Day, 2004, The Huygens Probe Was Released; - On A Ballistic Trajectory To Enter Titan's Atmosphere...



Jim Al-Khalili;
Narrator
Earl Maize;
Cassini-Huygens
Programme
Manager
Julie Webster;
Cassini
Spacecraft
Operations Team
Manager

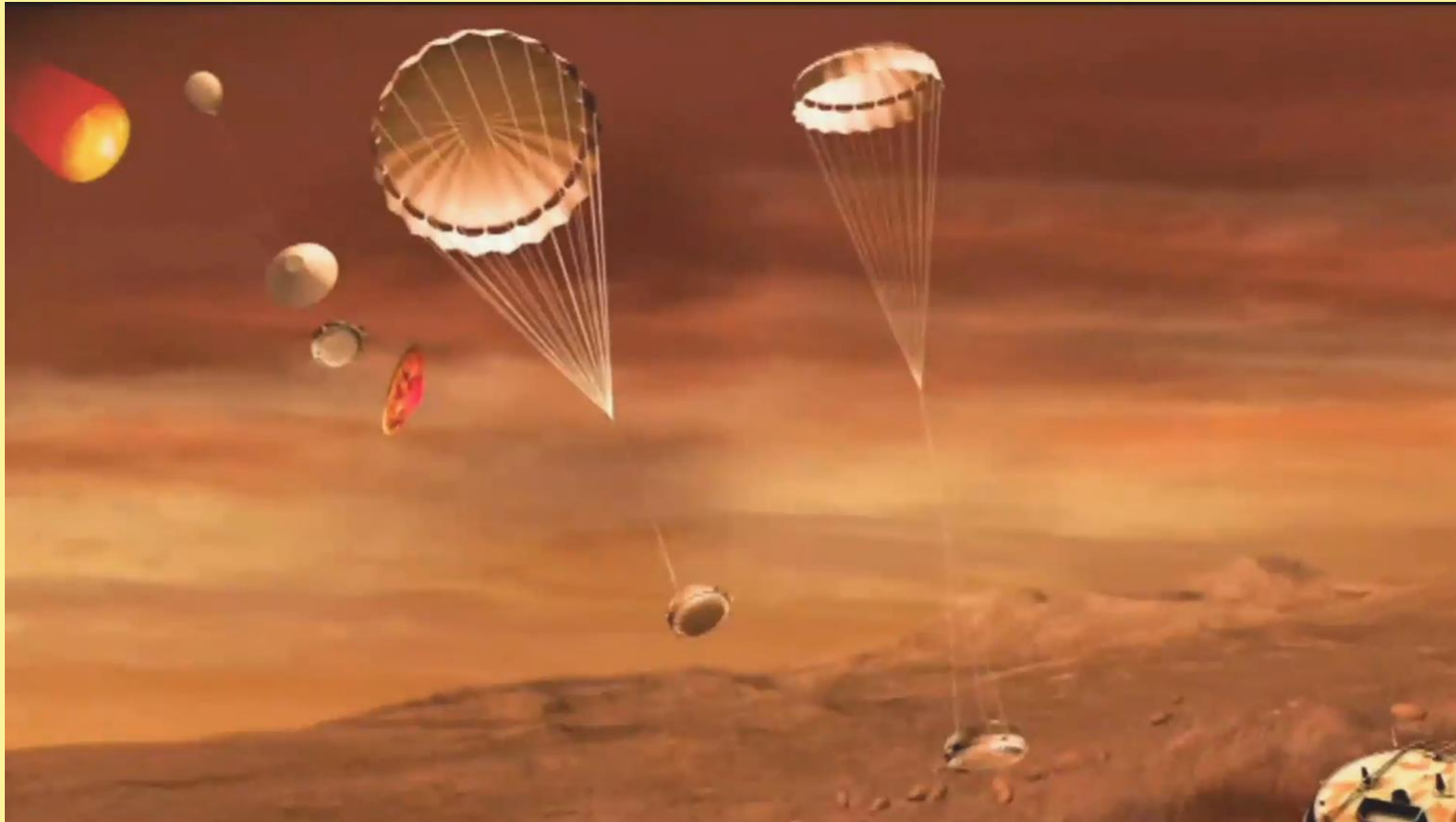
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After Coasting For 20 Days Huygens Entered Titan's Atmosphere At 6000 ms^{-1} , But 5 ms^{-1} Was Required To Land Safely On The Surface...



Manish Patel;
Planetary
Scientist
Earl Maize;
Cassini-Huygens
Programme
Manager
**Olivier
Witasse;**
ESA Planetary
Scientist
Jim Al-Khalili;
Narrator

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 - ...it landed on dry 'riverbed', & monitored -180°C environment for 72 mins



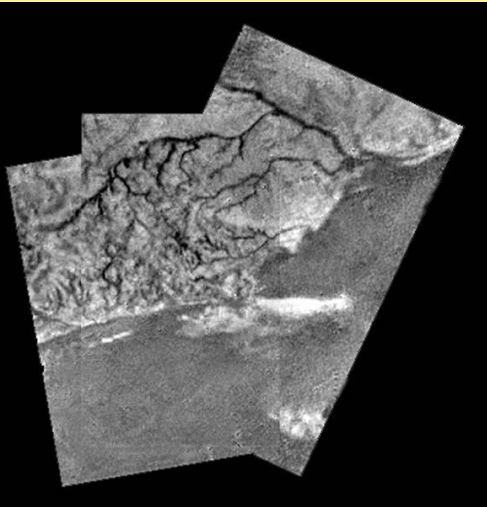
The Probe's Descent Took 150 Mins; - Longer Than Expected, But Huygen's Batteries Lasted 72 Mins On The Surface, Despite -180°C Temperature!



Elizabeth 'Zibi'
Turtle;
Principal
Investigator of
Dragonfly Mission
David
Harrington;
Narrator of final
descent
ANIMATION,
created by
University of
Arizona team
responsible for
the Descent
Imager / Spectral
Radiometer
(DISR)
instrument, which
captured the
images seen

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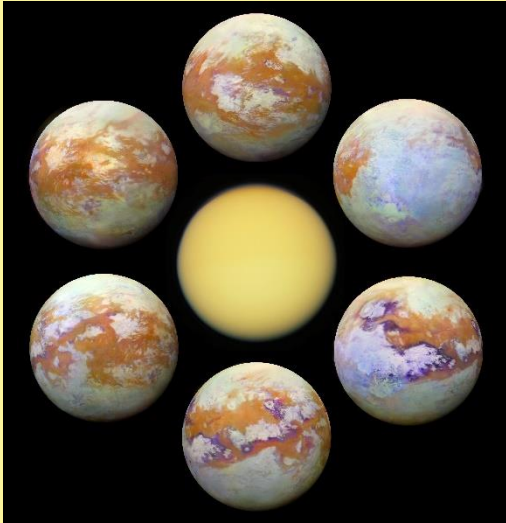
*The Legacy Of Liquid Was Clearly Evident On The Surface Of Titan,
Although The Probe Did Not Encounter ANY Liquid At The Landing Site...*



Michelle Dougherty;
Magnetometer
Team Leader,
Cassini Orbiter
John Zarnecki;
Principal
Investigator,
Huygens Surface
Science Package

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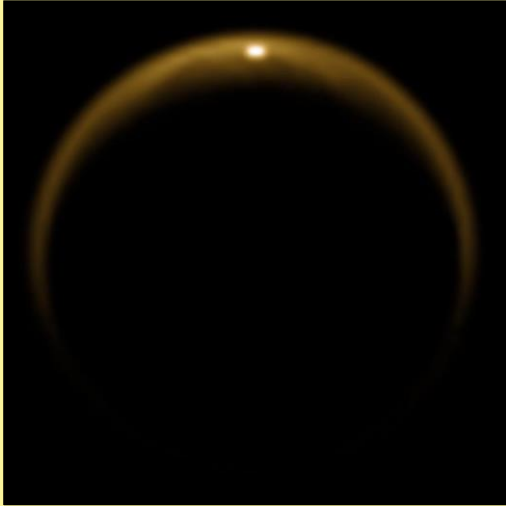
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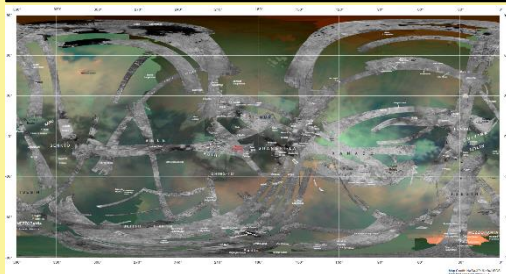
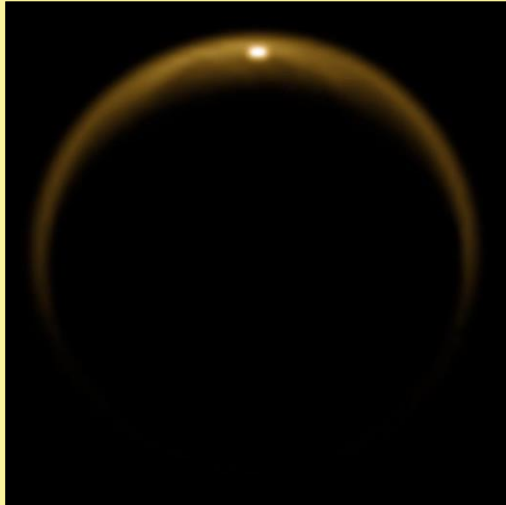


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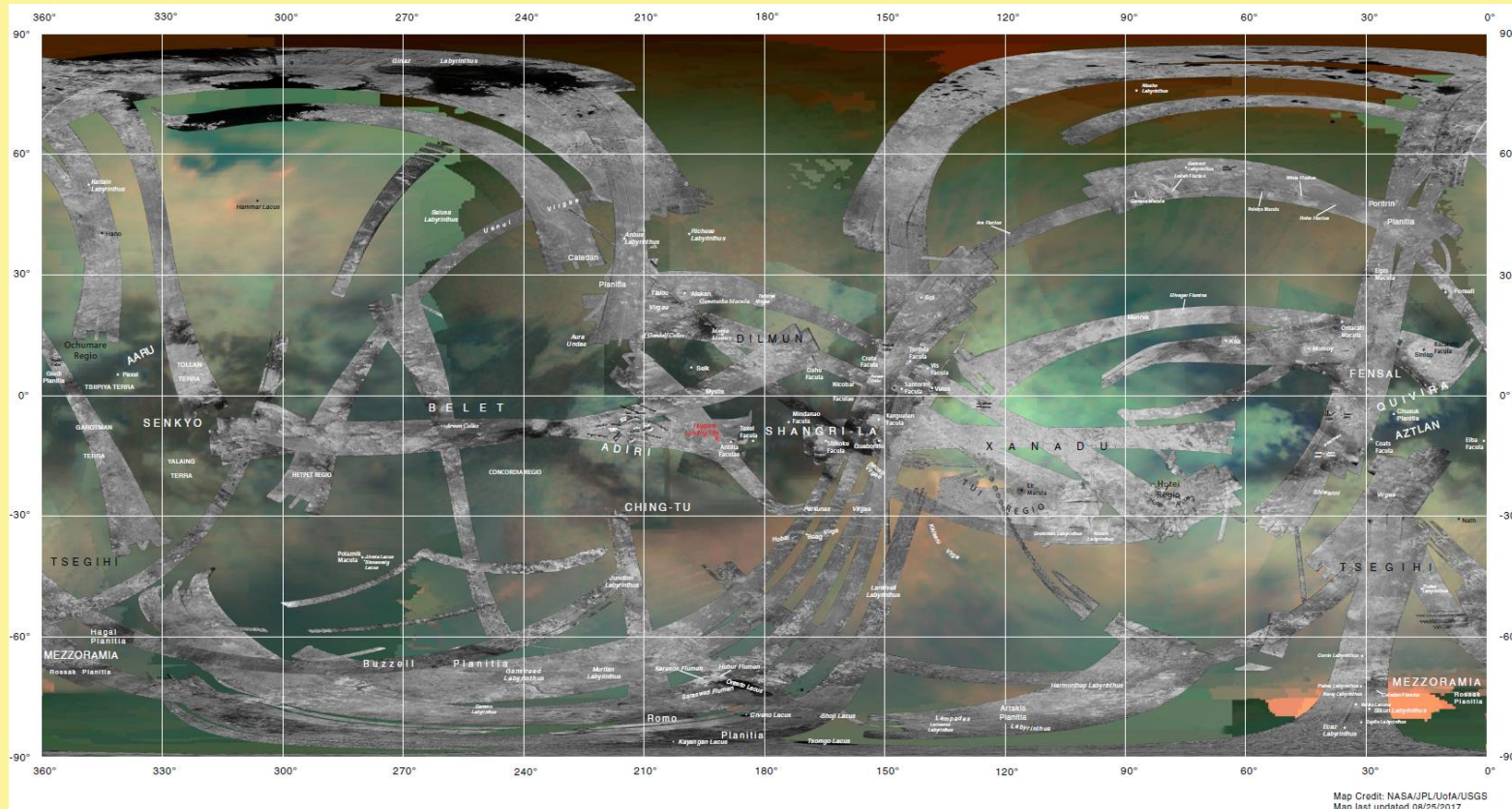
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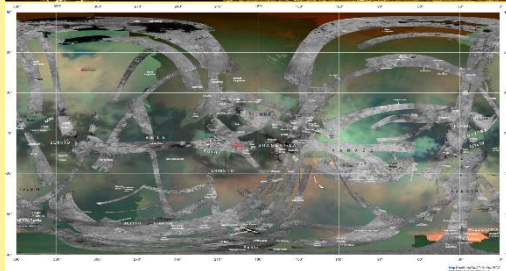
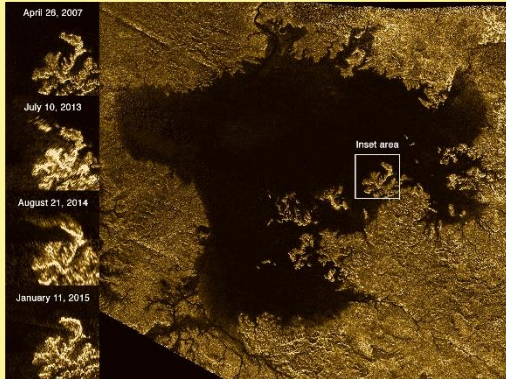
Cassini Radar Swaths (Monochrome) Overlaying VIMS Image Map (RGB); - Cylindrical Projection



Huygens
 Landing Site
 Lat; 10.6°S
 Long; 192.3°W
 Cassini radar
 Monochrome;
 $\lambda=2.18$ cm
 VIMS spectral
 mapping;
 R; $\lambda=5$ μ m
 G; $\lambda=2$ μ m
 B; $\lambda=1.3$ μ m
 The very 'radar
 dark' patches in
 NW are the
 three major
 liquid 'seas' on
 Titan, viz;
 Punga Mare,
 Kraken Mare &
 Ligeia Mare

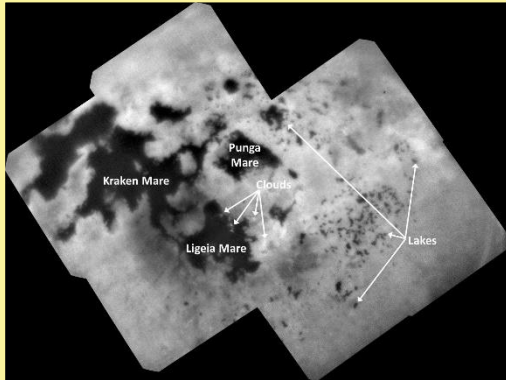
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...and RADAR has observed transient features in maria; e.g. Ligeia Mare...

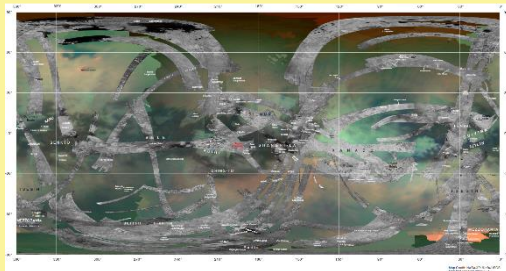


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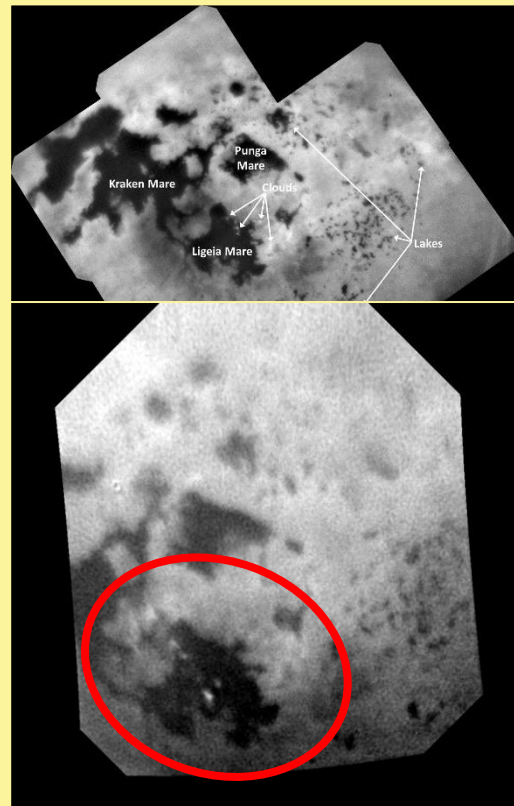


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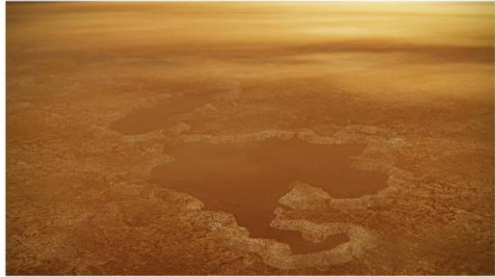
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...and has observed relatively rare cloud development; Note Ligeia Mare...

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Sep 9, 2019

New Models Suggest Titan Lakes Are Explosion Craters



This artist's concept of a lake at the north pole of Saturn's moon Titan illustrates raised rims and rampartlike features such as those seen by NASA's Cassini spacecraft around the moon's Winnipeg Lacus. Credit: NASA/JPL-Caltech

Using radar data from NASA's Cassini spacecraft, recently published research presents a new scenario to explain why some methane-filled lakes on Saturn's moon Titan are surrounded by steep rims that reach hundreds of feet high. The models suggest that explosions of warming nitrogen created basins in the moon's crust.

Titan is the only planetary body in our solar system other than Earth known to have stable liquid on its surface. But instead of water raining down from clouds and filling lakes and seas as on Earth, on Titan it's methane and ethane — hydrocarbons that we think of as gases but that behave as liquids in Titan's frigid climate.

Most existing models that lay out the origin of Titan's lakes show liquid methane dissolving the moon's bedrock of ice and solid organic compounds, carving reservoirs that fill with the liquid. This may be the origin of a type of lake on Titan that has sharp boundaries. On Earth, bodies of water that formed similarly, by dissolving surrounding limestone, are known as karstic lakes.

The new, alternative models for some of the smaller lakes (tens of miles across) turns that theory upside down: it proposes pockets of liquid nitrogen in Titan's crust warmed, turning into explosive gas that blew out craters, which then filled with liquid methane. The new theory explains why some of the smaller lakes near Titan's north pole, like Winnipeg Lacus, appear in radar imaging to have very steep rims that tower above sea level — rims difficult to explain with the karstic model.

The radar data were gathered by the Cassini Saturn Orbiter — a mission managed by NASA's Jet Propulsion Laboratory in Pasadena, California — during its last close flyby of Titan, as the spacecraft prepared for its final plunge into Saturn's atmosphere two years ago. An international team of scientists led by Giuseppe Mitri of Italy's G. d'Annunzio University became convinced that the karstic model didn't jibe with what they saw in these new images.

"The rim goes up, and the karst process works in the opposite way," Mitri said. "We were not finding any explanation that fit with a karstic lake basin. In reality, the morphology was more consistent with an explosion crater, where the rim is formed by the ejected material from the crater interior. It's totally a different process."

The work, published Sept. 9 in *Nature Geosciences*, meshes with other Titan climate models showing the moon may be warm compared to how it was in earlier Titan "ice ages."

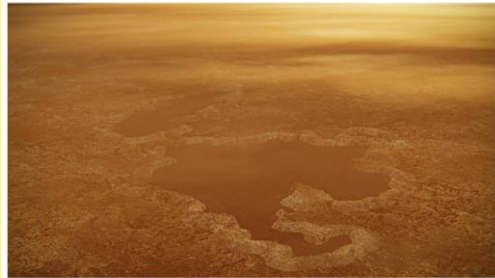
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- Many questions remain

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New Models Suggest Titan Lakes Are Explosion Craters



This artist's concept of a site at the north pole of Saturn's moon Titan illustrates relief and ripples features such as those seen by NASA's Cassini spacecraft around the moon's Winnipeg Lacus. Credit: NASA/JPL-Caltech

Using radar data from NASA's Cassini spacecraft, recently published research presents a new scenario to explain why some methane-filled lakes on Saturn's moon Titan are surrounded by steep rims that reach hundreds of feet high. The models suggest that explosions of warming nitrogen created basins in the moon's crust.

Titan is the only planetary body in our solar system other than Earth known to have stable liquid on its surface. But instead of water raining down from clouds and filling lakes and seas as on Earth, on Titan it's methane and ethane — hydrocarbons that we think of as gases but that behave as liquids in Titan's frigid climate.

Most existing models that lay out the origin of Titan's lakes show liquid methane dissolving the moon's bedrock of ice and solid organic compounds, carving reservoirs that fill with the liquid. This may be the origin of a type of lake on Titan that has sharp boundaries. On Earth, bodies of water that formed similarly, by dissolving surrounding limestone, are known as karstic lakes.

The new, alternative models for some of the smaller lakes (tens of miles across) turns that theory upside down: it proposes pockets of liquid nitrogen in Titan's crust warmed, turning into explosive gas that blew out craters, which then filled with liquid methane. The new theory explains why some of the smaller lakes near Titan's north pole, like Winnipeg Lacus, appear in radar imaging to have very steep rims that tower above sea level — rims difficult to explain with the karstic model.

The radar data were gathered by the Cassini Saturn Orbiter — a mission managed by NASA's Jet Propulsion Laboratory in Pasadena, California — during its last close flyby of Titan, as the spacecraft prepared for its final plunge into Saturn's atmosphere two years ago. An international team of scientists led by Giuseppe Mitri of Italy's G. d'Annunzio University became convinced that the karstic model didn't jibe with what they saw in these new images.

"The rim goes up, and the karst process works in the opposite way," Mitri said. "We were not finding any explanation that fit with a karstic lake basin. In reality, the morphology was more consistent with an explosion crater, where the rim is formed by the ejected material from the crater interior. It's totally a different process."

The work, published Sept. 9 in Nature Geosciences, meshes with other Titan climate models showing the moon may be warm compared to how it was in earlier Titan "ice ages."

- Cassini-Huygens was instrumented to penetrate atmosphere
 - CCD-based wide- & narrow-angle cameras, with 938 nm NIR filter (ISS)
 - Cassini (Ku-band, synthetic aperture; $\lambda = 2.18$ cm) radar (RADAR)
 - Visible and Infrared Mapping Spectrometer (VIMS) + Huygens probe
- Huygens departed Cassini on it's 3rd orbit; Xmas Day, 2004
...and entered atmosphere on 14-Jan-2005, using heatshield and 'chutes
 - 2.5 hr descent went well; - MOST data sent to Cassini, despite glitch...
...it landed on dry 'riverbed', & monitored -180°C environment for 72 mins
- DISR imaged liquid legacy, but probe encountered no liquid...
 - But atmosphere transmits 1.3-5 μm , so VIMS can see surface detail...
...in 2009 VIMS caught a 'glint' of sun reflected from Northern lake surface...
 - RADAR discriminates liquids, which VIMS says are ethane & methane
...and RADAR has observed transient features in maria; e.g. Ligeia Mare...
 - NIR-filtered ISS penetrates atmosphere to map seas & lakes in North...
...and has observed relatively rare cloud development; Note Ligeia Mare...
- Many questions remain; - **more 'Ground Truth' is required...**

Credits 1; - Speakers to camera (in order of first appearance)

'critical science objective of Voyager'; BBC4 Storyville strand; "The Farthest: Voyager's Interstellar Journey"; <https://www.bbc.co.uk/programmes/b09gvnty>; Last shown 2019

Carl Saga (Voiceover) Astrophysicist, Astronomer, Cosmologist, Scientist

John Casani Voyager Project Manager

Jim Bell Author of 'Interstellar Age'

'camera were defeated by 'smog'; BBC4 Storyville strand; "The Farthest: Voyager's Interstellar Journey"; <https://www.bbc.co.uk/programmes/b09gvnty>; Last shown 2019

Jim Bell Author of 'Interstellar Age'

Larry Soderblom Voyager Imaging Science Team (Astrogeologist)

'Cassini-Huygens was instrumented to penetrate atmosphere'; BBC Horizon strand; "Saturn: Lord of the Rings"; <https://www.bbc.co.uk/programmes/p00607tq>; 2004-5 Season

Larry Esposito Principal Investigator, Cassini UVIS Instrument

Jack Fortune (Narrator) Narrator; Horizon; 'Lord Of The Rings'

'ISS'; BBC Horizon strand; "Saturn: Lord of the Rings";
<https://www.bbc.co.uk/programmes/p00607tq>; 2004-5 Season

Jack Fortune Narrator ; Horizon; 'Lord Of The Rings'

Carolyn Porco ISS Team Leader, Cassini

*Credits 2; - Speakers to camera
(in order of first appearance)*

'RADAR'; BBC Horizon strand; "Saturn: Lord of the Rings";

<https://www.bbc.co.uk/programmes/p00607tg>; 2004-5 Season

Jack Fortune Narrator ; Horizon; ‘Lord Of The Rings’

'VIMS'; BBC Horizon strand; "Saturn: Lord of the Rings";

<https://www.bbc.co.uk/programmes/p00607tg>; 2004-5 Season

Jack Fortune	Narrator ; Horizon; 'Lord Of The Rings'
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Bob Brown VIMS Team Leader, Cassini

'VIMS'; BBC Horizon strand; "Saturn: Lord of the Rings";

<https://www.bbc.co.uk/programmes/p00607tg>; 2004-5 Season

Jack Fortune Narrator ; Horizon; ‘Lord Of The Rings’

John Zarnecki Principal Investigator, Huygens Surface Science Package

'Huygens departed Cassini'; Yesterday Channel; "Secrets of the Solar System; Episode 7; - Saturn"; <https://uktvplay.uktv.co.uk/shows/secrets-of-the-solar-system/watch-online/6166352070001>; 2019

Animation of Huygens's trajectory from 25 December 2004 to 14 January 2005,

https://upload.wikimedia.org/wikipedia/commons/7/72/Animation_of_Huygens_trajectory.gif

Jim Al-Khalili (Narrator)	Narrator; Secrets Of The Solar System; 'Saturn'
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Earl Maize

Julie Webster Cassini Spacecraft Operations Team Manager

Credits 3: - Speakers to camera (in order of first appearance)

'heatshields and 'chutes'; Yesterday Channel; "Secrets of the Solar System; Episode 7; - Saturn";
<https://uktvplay.uktv.co.uk/shows/secrets-of-the-solar-system/watch-online/6166352070001>; 2019

Manish Patel

Planetary Scientist

Earl Maize

Cassini-Huygens Program Manager

Olivier Witasse

ESA Planetary Scientist

Jim Al-Khalili (Narrator)

Narrator; Secrets Of The Solar System; 'Saturn'

'descent' ; Yesterday Channel; "Secrets of the Solar System; Episode 7; - Saturn";
<https://uktvplay.uktv.co.uk/shows/secrets-of-the-solar-system/watch-online/6166352070001>; 2019

Erich Karkoschka, Lunar and Planetary Lab, DISR team, University of Arizona;
<https://sci.esa.int/web/cassini-huygens/-/39218-huygens-descent-to-titan-surface>; 2015

Elizabeth 'Zibi' Turtle

Planetary Scientist

David Harrington (Narrator)

University of Arizona Huygens Lander final descent animation

'encountered no liquid'; BBC, The Sky At Night; "Unveiling Titan";
<https://www.bbc.co.uk/iplayer/episode/p00f8xfz/the-sky-at-night-unveiling-titan>; 7 Feb 2005

Michelle Dougherty

Magnetometer Team Leader, Cassini

John Zarnecki

Principal Investigator, Huygens Surface Science Package

Patrick Moore

(Silent observer)

Thank You For Your Attention!

Any questions?